



# Studi e missioni iperspettrali nel VIS-SWIR : contributo della missione PRISMA allo studio di fenomeni geofisici

Istituto Nazionale di Geofisica e Vulcanologia:

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Workshop Data Exploitation della missione PRISMA,  
precursore delle missioni iperspettrali nazionali



Roma 11-3 Marzo 2017

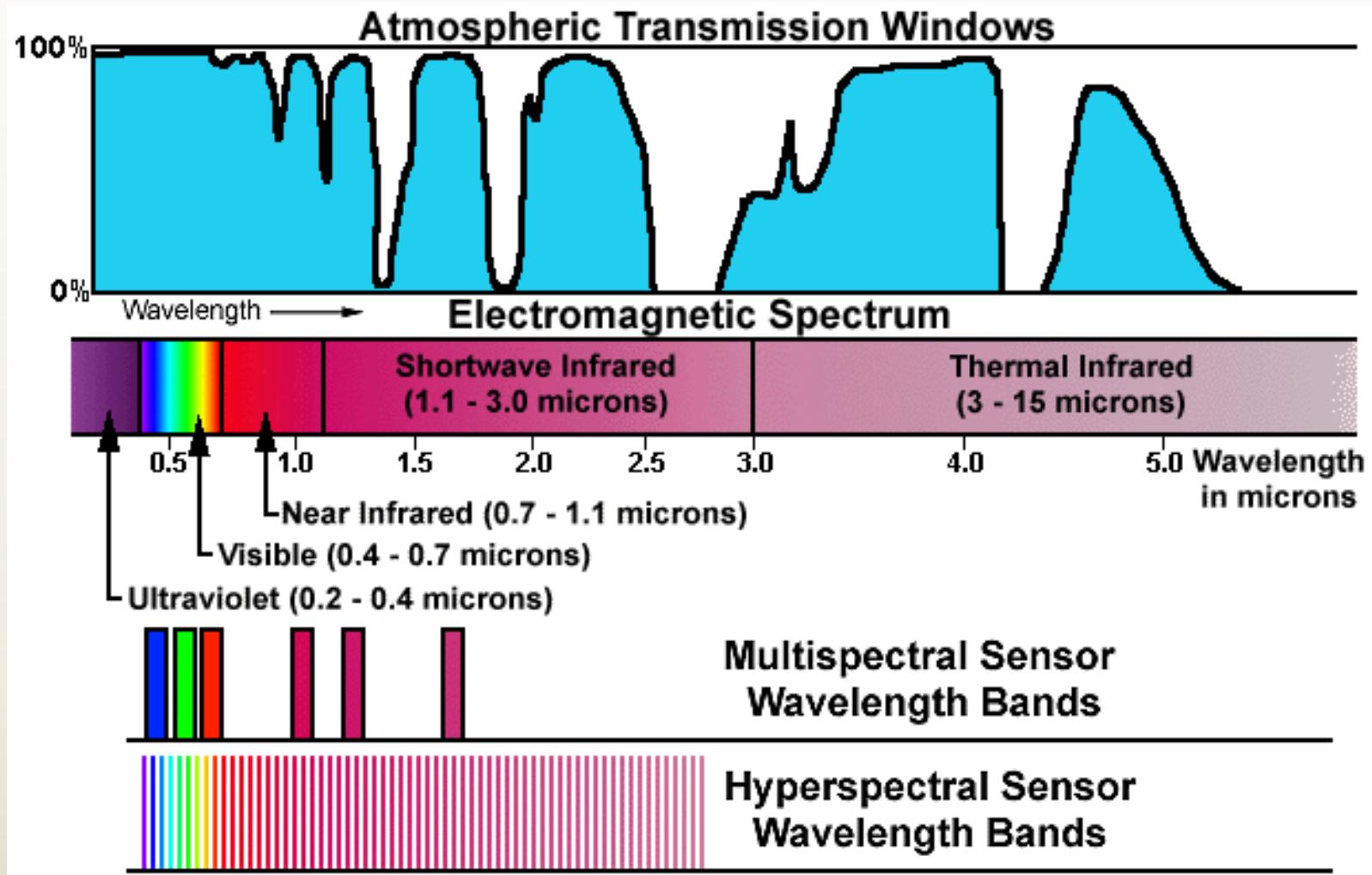




# SUMMARY

1. Image spectrometers :satellite missions
2. Hyperspectral images data to monitor  
Geophysical hazards
1. ASI-AGI Project applications
2. Processing chains concept
3. Some results
4. Conclusion

# IMAGE SPECTROMETER (HYPERSPECTRAL SENSORS SPECTRAL RANGES VIS-IR)



## Hyperspectral Missions – Launch and Lifetime

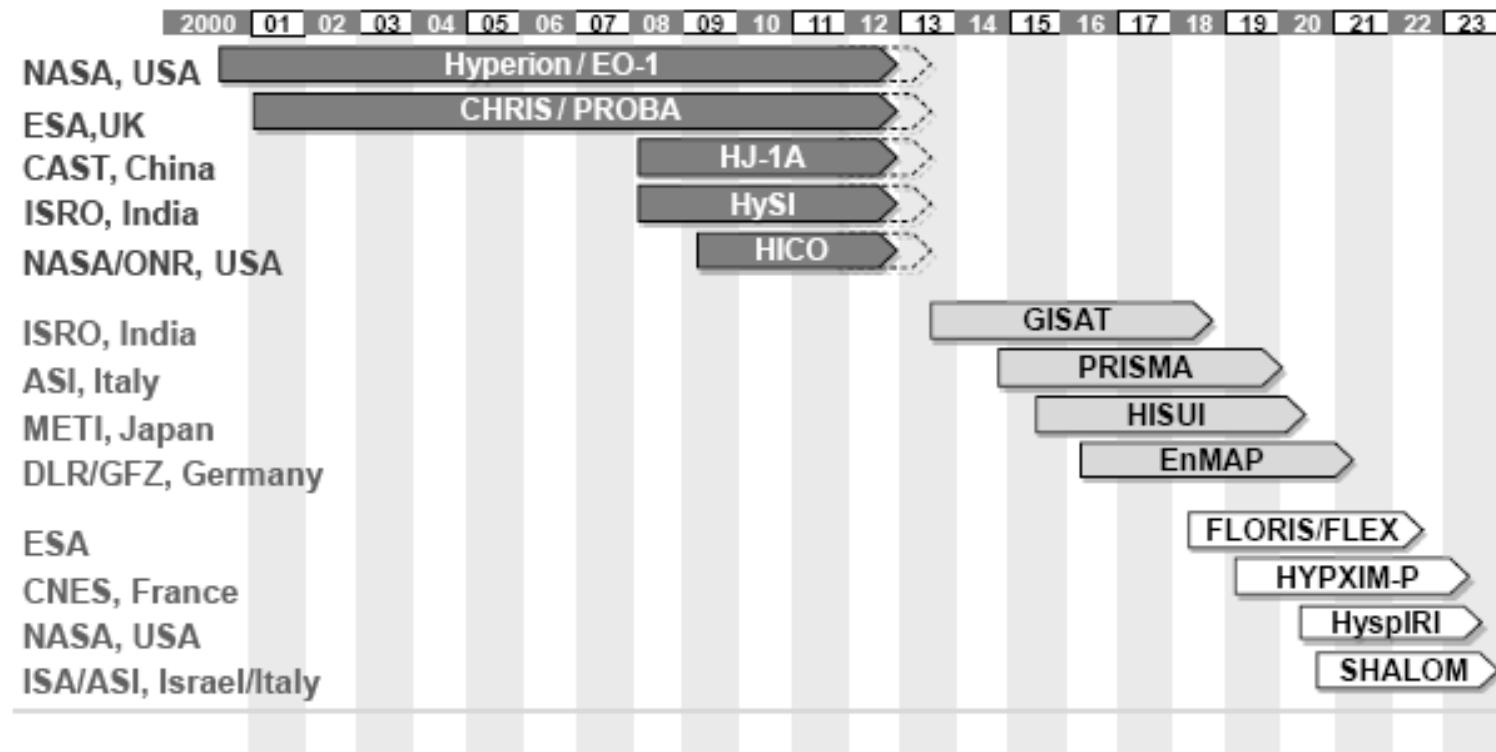
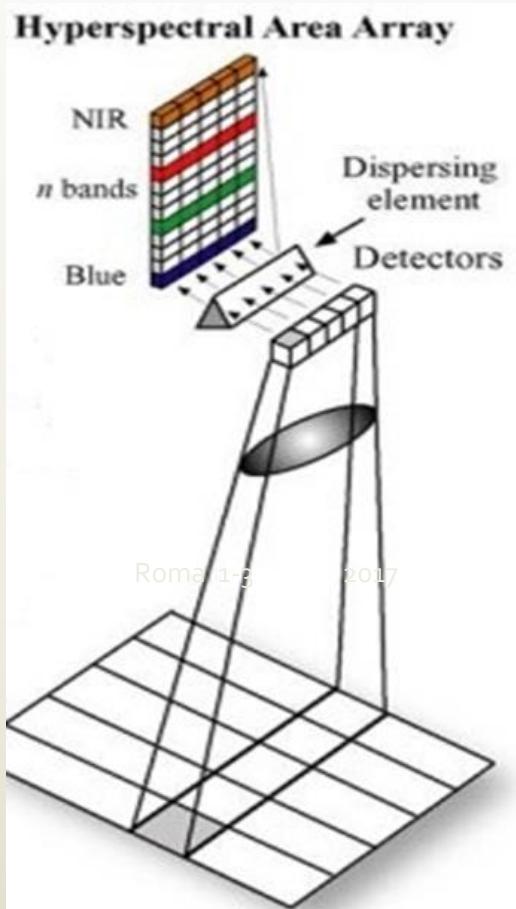


Figure 2: Spaceborne imaging spectroscopy mission launch dates and lifetimes (█ missions currently in operation, □ missions under construction, ▨ missions in a planning stage).

# Current and Planned Civilian Satellite Missions

	VIS-NIR	SWIR	TIR
Hyperion EO-1 (USA 2000 - )			30 m
Chris/Proba EU 2001 - )			17/34 m
HySI (India 2008 - )			500 m
HJ-1A (China, 2008 - )			100 m
MimSAT (South Africa 2010?)			30 m
PRISMA (Italy 2012 )			30 m
EnMAP Germany 2013 )			30 m
Hyper/Multi Japan 2013 )			?
HyspIRI (USA 2014? - )			60 m

The PRISMA Payload is an Electro-Optical instrument based on a **pushbroom scanning** technique, consisting of a high spectral resolution imaging spectrometer. It operates in the spectral range 0.4-2.5  $\mu\text{m}$  and is optically integrated with a medium resolution Panchromatic camera operating in the spectral range 0.4-0.7  $\mu\text{m}$ .

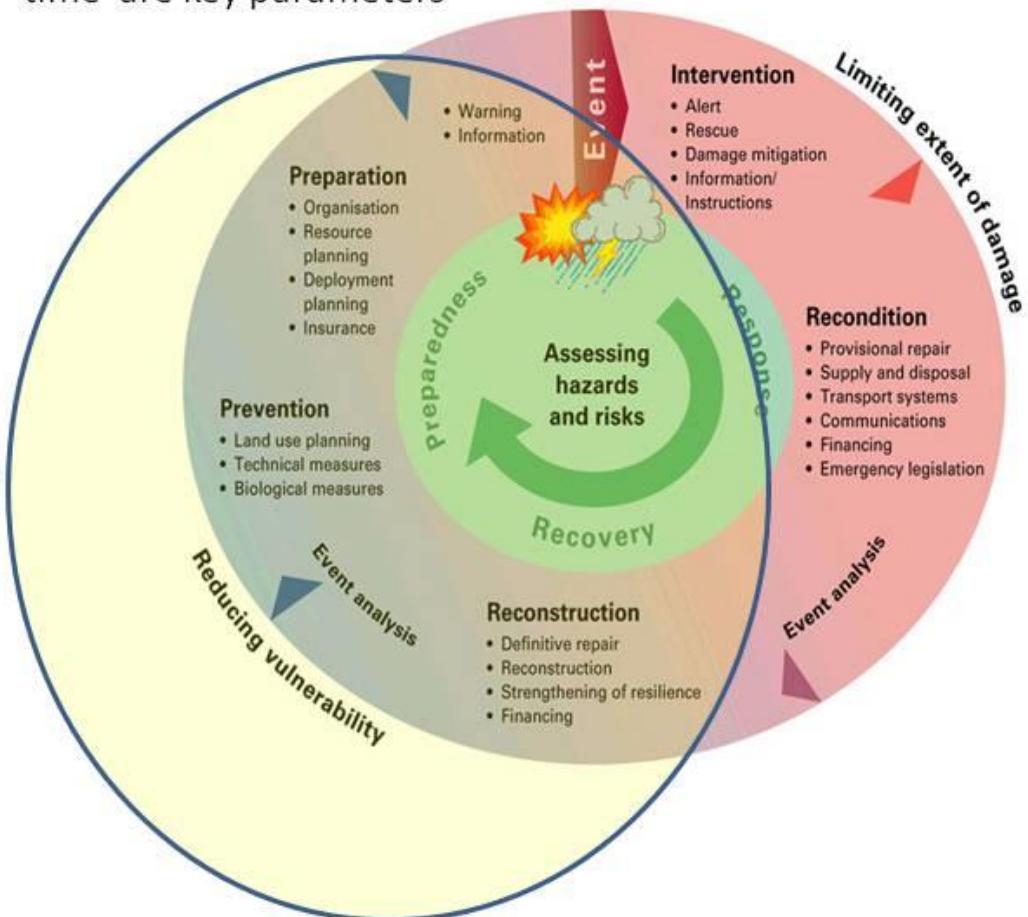


Instrument Main Characteristics	
Swath / FOV	30 km / 2.45°
GSD	<ul style="list-style-type: none"> <li>Hyperspectral: 30 m</li> <li>PAN: 5 m</li> </ul>
Spatial Pixels	Hyperspectral: 1000 PAN: 6000
Spectral Range	VNIR: 400 – 1010 nm SWIR: 920 – 2505 nm
Spectral Resolution	$\leq 12 \text{ nm}$
Spectral Bands	VNIR: 66 SWIR: 171
Radiometric Quantization	12 bit
VNIR SNR	$> 200:1$ on 400 – 1000 nm $> 500:1$ @ 650 nm
SWIR SNR	$> 200:1$ on 1000 – 1750 nm $> 400:1$ @ 1550 nm $> 100:1$ on 1950 – 2350 nm $> 200:1$ @ 2100 nm
PAN SNR	$> 240:1$
Absolute Radiometric Accuracy	Better than 5%

Courtesy of ASI



Hyperspectral data could provide important information in all the Disaster management Phases with a specific contribution to: **Preparation, Mitigation and Recovery Phases**. Spectral, spatial resolution and satellite revisit time are key parameters

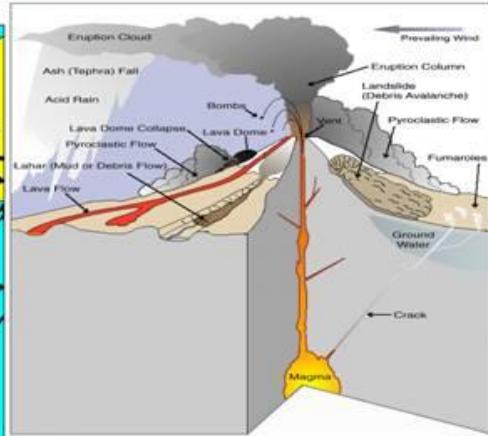
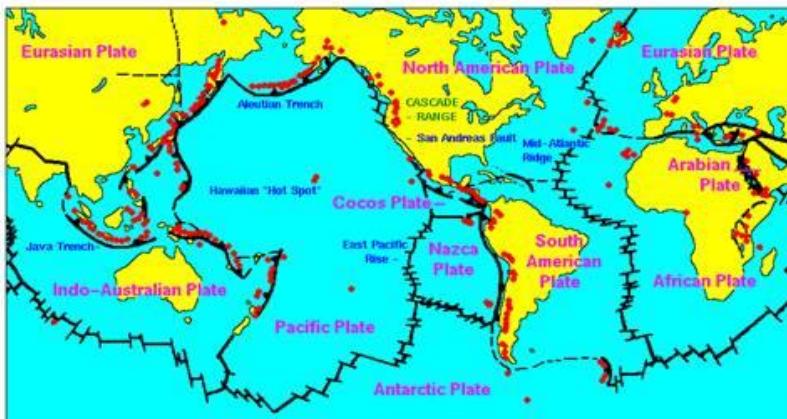


PRISMA  
APPLICATIONS





## PRISMA PRODUCTS TO MONITOR VOLCANOES



RISK PHASE	PRISMA PRODUCTS	SPECIFICATIONS
Prevention Phase	Lava product mapping: mineral analysis Vegetation growth and weathering Ash deposits	<b>Full spectral range</b> Data fusion with panchromatic
	Degassing plumes analysis (H <sub>2</sub> O, CO <sub>2</sub> ), volcanic aerosols	<b>Full spectral range</b> Water and CO <sub>2</sub> bands mandatory
Crisis Phase	active lava flow Thermal Flux, Effusion Rate	SWIR range multi-channel approach, avoid saturation
	Mapping the development of new Lava flows	SWIR channels, data fusion with panchromatic
Post Crisis Phase	Volcanic products impact on soil and vegetation	<b>Full spectrum, integration with ground measurements</b>
	Update of geological maps and risk maps	<b>Full spectrum, data fusion with panchromatic, geometric corrections</b>

## INGV infrastructure and services based on EO data in support of the Civil Protection Department (DPC) also considered for COPERNICUS services

1991-2000

- Partecipazione alla campana MAC EUROPE 91
- Progetti di ricerca DPC
- Sviluppo procedure per l'analisi EO per applicazioni vulcanologiche e sismiche dati OTTICI e SAR
- Progetti MADVIEWS ed EMPEDOCLE

2001-2006

- Acquisizione dati EO in modo sistematico (sistemi di acquisizione RT)
- Sviluppo procedure operative su requisiti forniti da DPC
- Monitoraggio con dati EO viene introdotto nelle convenzioni DPC-INGV

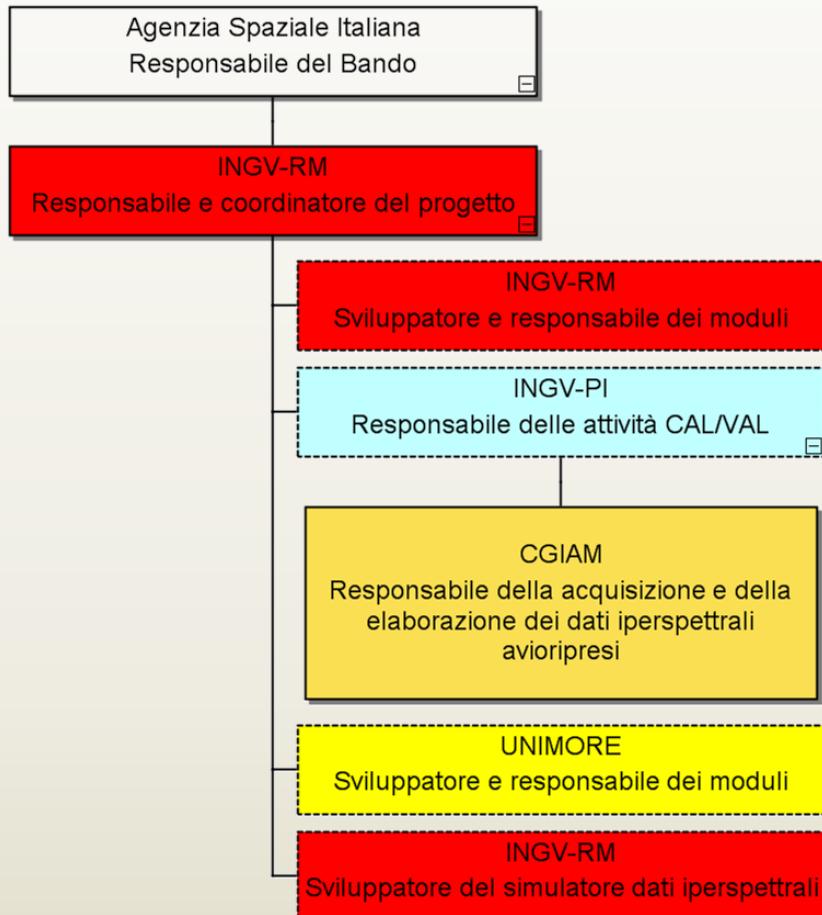
2004-2011

- 2004 -2011 Sviluppo PROGETTI PILOTA ASI per il supporto al rischio Vulcanico e Sismico
- Requisiti dei prodotti sono definiti del DPC

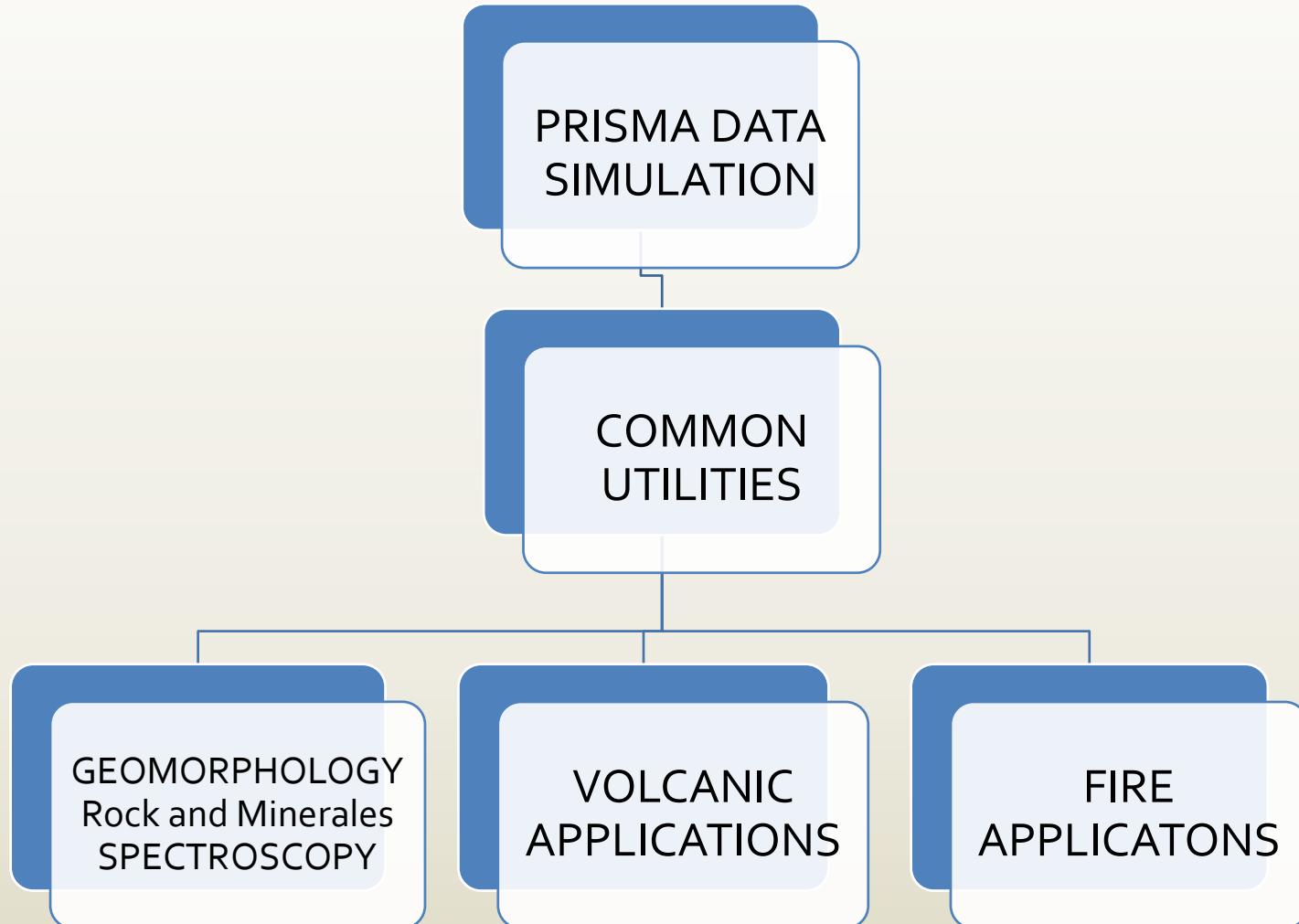
2006-2014

- INGV ha partecipato allo sviluppo del sistema GMES-COPERNICUS guidando il settore VULCANI e TERREMOTI per i servizi del CORE SERVICE EMERGENCY

# ASI-AGI: ANALISI SISTEMI IPERSPETRALI PER LE APPLICAZIONI GEOFISICHE INTEGRATE



# PRISMA GEOPHYSICAL APPLICATION PRODUCTS (ASI-AGI)



# PRODUCTS DEVELOPMENT APPROACH

## COMMON UTILITIES

Atmospheric correction

Pan Sharpening

Topographic Correction



## ADDED VALUES VOLCANIC PRODUCTS

Volcanic products  
Classification

Lava Thermal analysis

Volcanic gases analysis

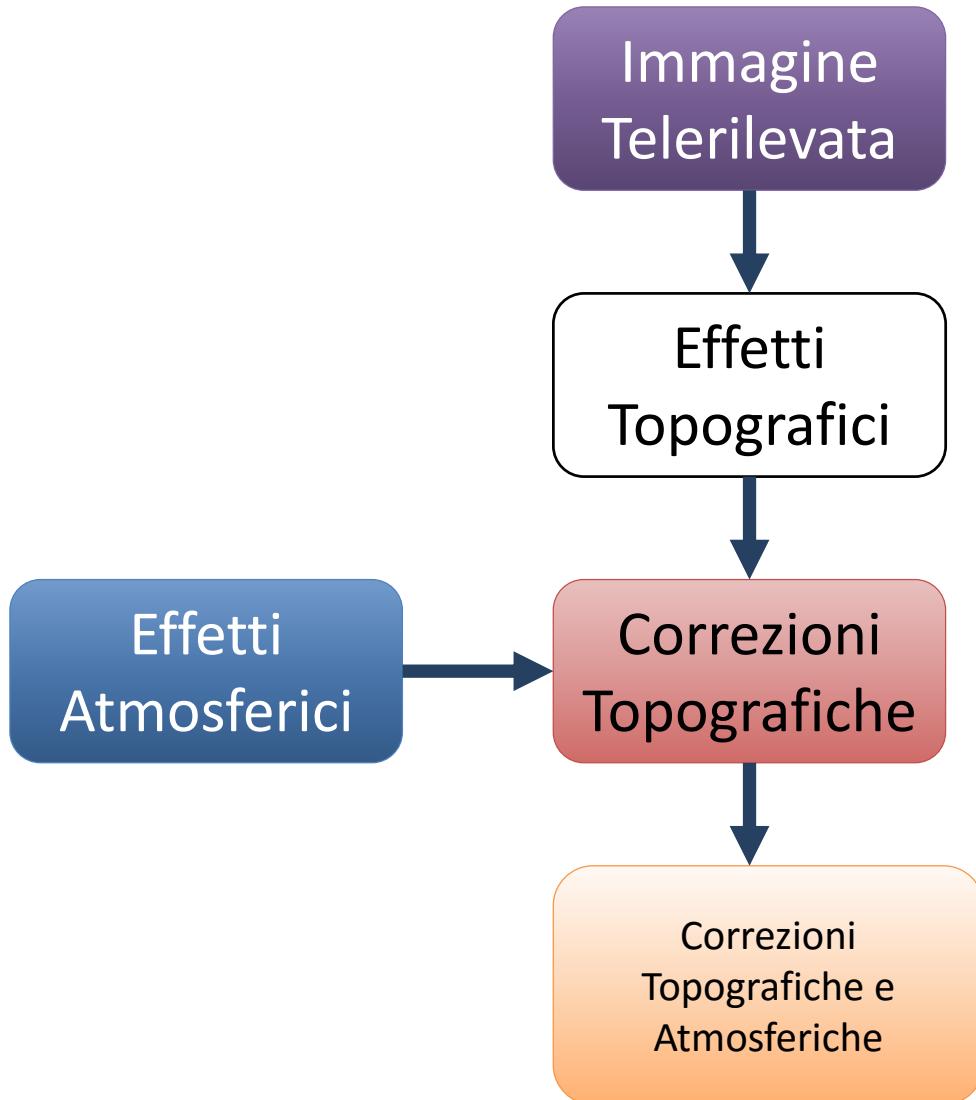


## CAL / VAL ACTIVITIES

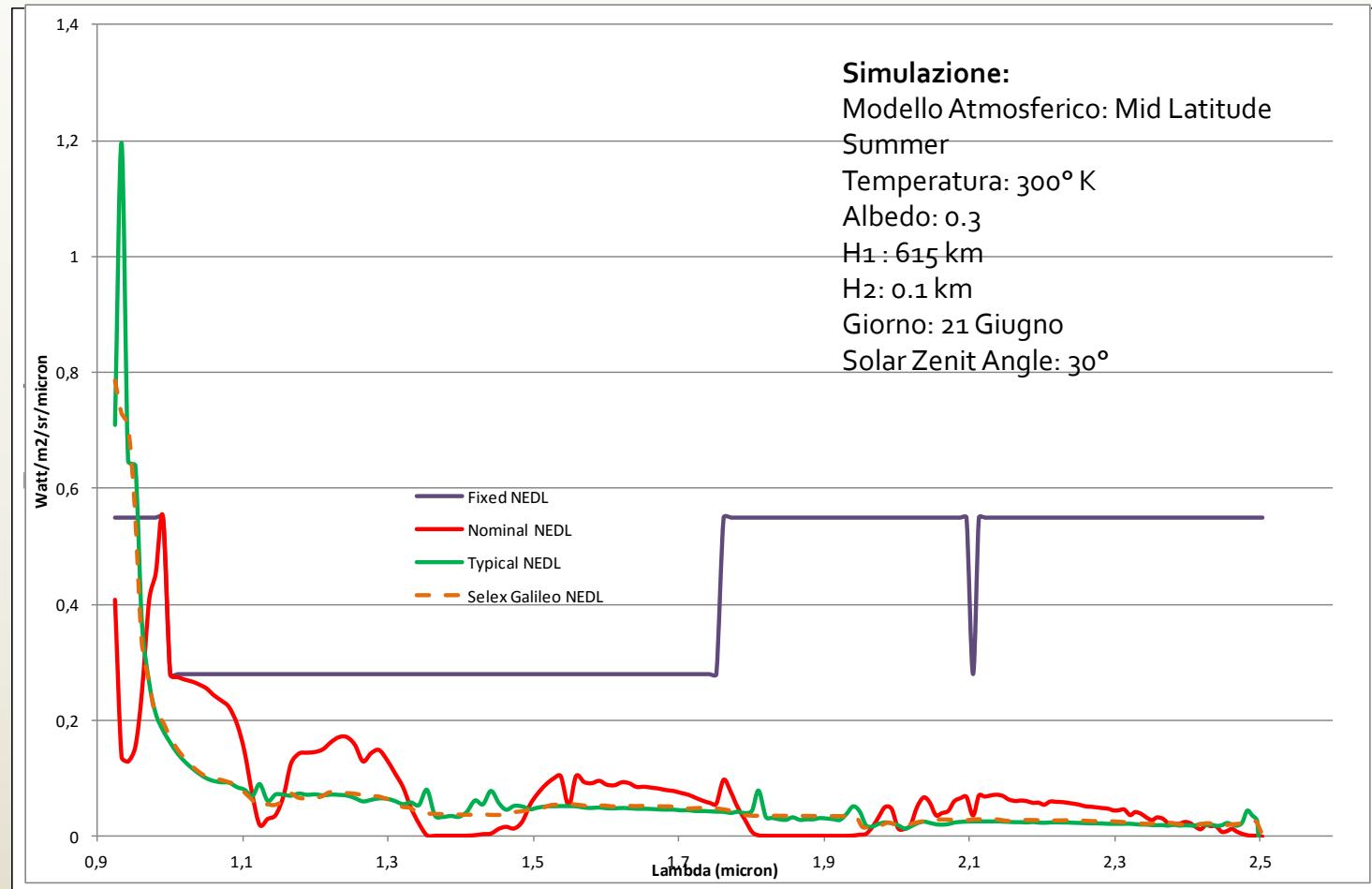
PERMANENT TEST AREAS

MT. Etna spectral library

# Correzione Topografiche e Atmosferiche

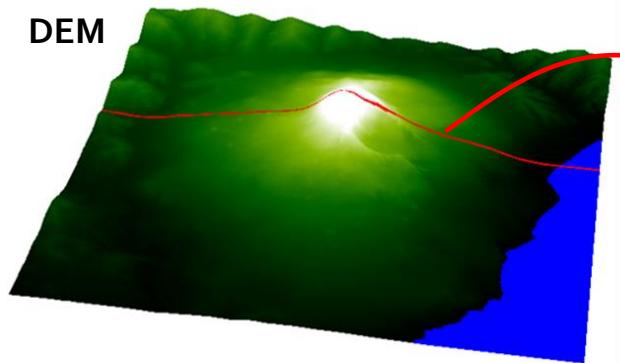


# PRISMA SIMULATED TRANSMITTANCE RADIANCE and NeDL

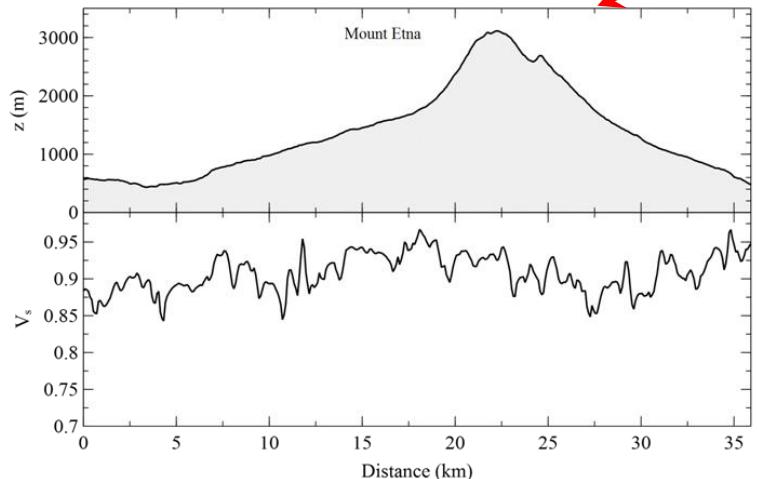
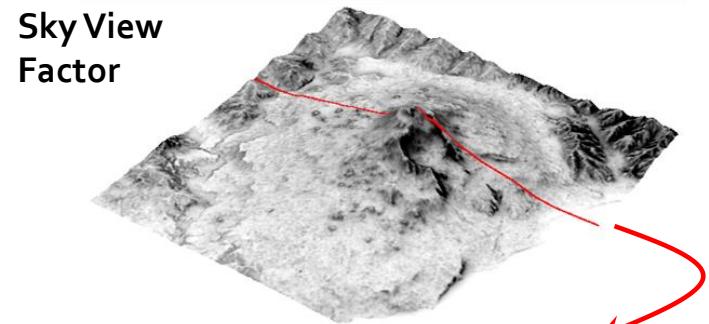


## WP 2.1 Correzione topografica: "Screenshoots"

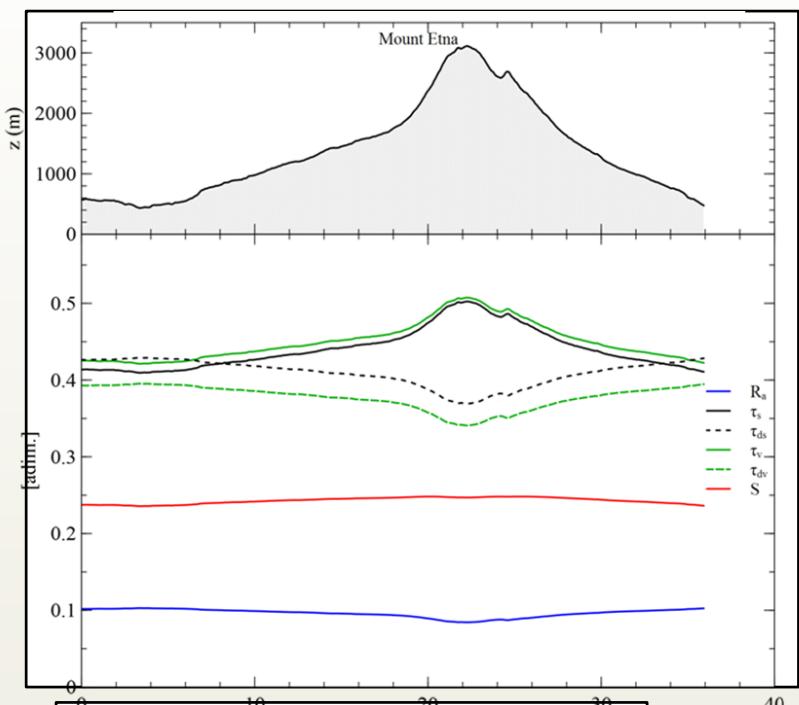
DEM



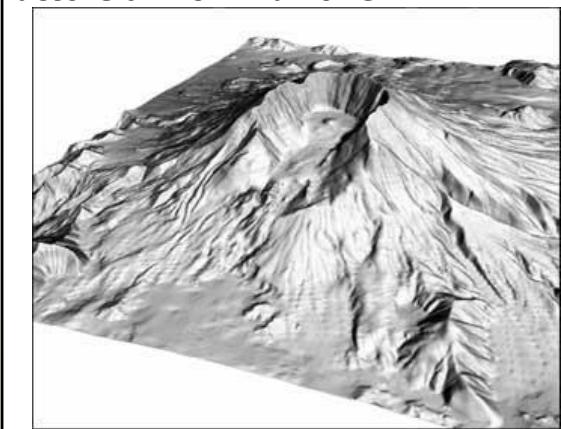
Sky View Factor



Termini di correzione atmosferica

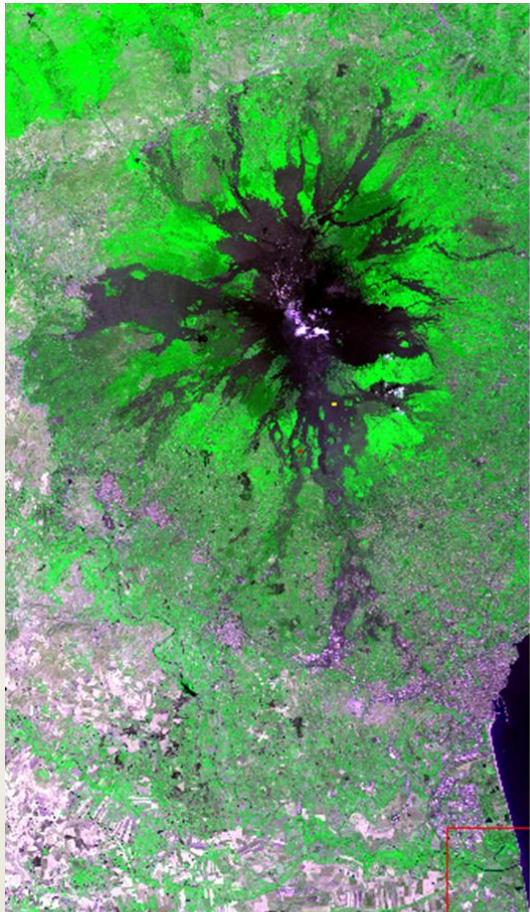


Fattore di illuminazione



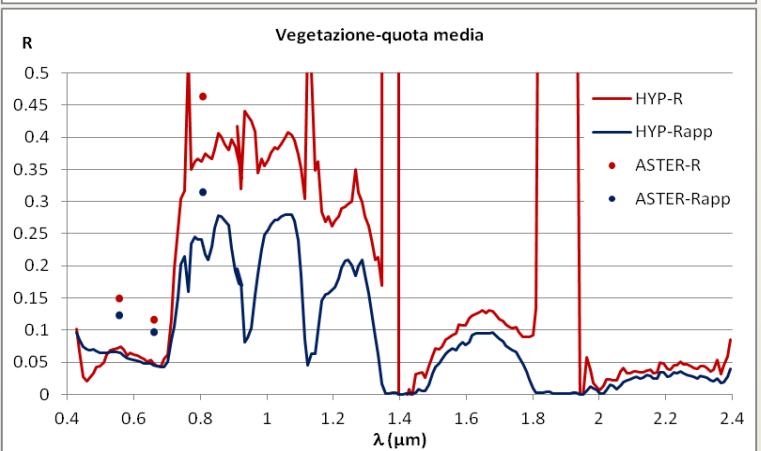
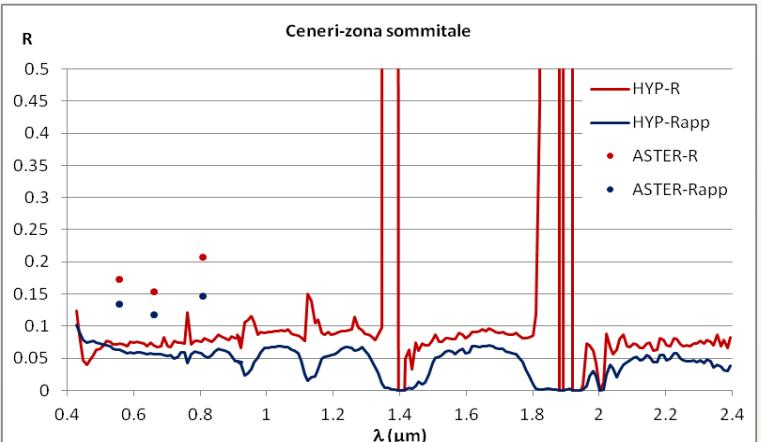
## ASTER

25/06/2012



## HYPERION

26/06/2012



Riflettanze apparente e corrette dalle TAC  
estratte da due aree comuni sulle immagini  
ASTER e HYPERION  
Problemi di Calibrazione!

topographic corrections: Some results

## CARATTERISTICHE DEL SENSORE PRISMA DI INTERESSE:

### Sensore Iperspettrale (HS)

0.4 – 2.5  $\mu\text{m}$

(risoluzione spettrale: 10 nm)

210 canali

Risoluzione spaziale:

30m

### Sensore Pancromatico (PAN)

0.4 – 0.75  $\mu\text{m}$

1 canale

Risoluzione spaziale:

5m



**SCOPO:** “Fusione” delle due immagini acquisite in modo da ottenere una nuova immagine sintetica con la risoluzione spaziale del sensore PAN e la risoluzione spettrale del sensore HS



Pan sharpening

# DESCRIZIONE DELLA PROCEDURA

Immagine HS PRISMA – Bassa  
risoluzione spaziale –  
*Livello L2*

Immagine PAN PRISMA – Alta  
risoluzione spaziale –  
*Livello L2*

Eventuale co-registrazione delle immagini

Definizione della destinazione d'uso  
delle immagini sintetiche prodotte

Scelta dell'algoritmo di  
pansharpening idoneo

Implementazione dell'algoritmo

Immagine Sintetica pansharpened

Valutazione della qualità della  
fusione

PC

Wavelet

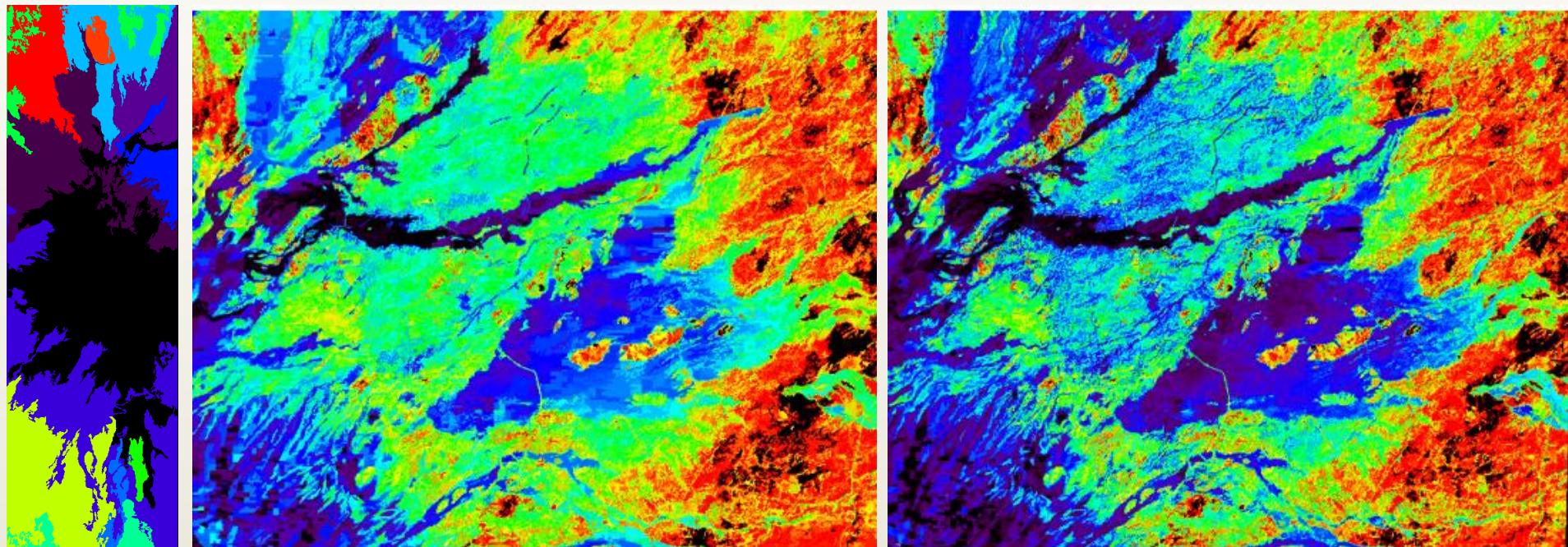
HCS

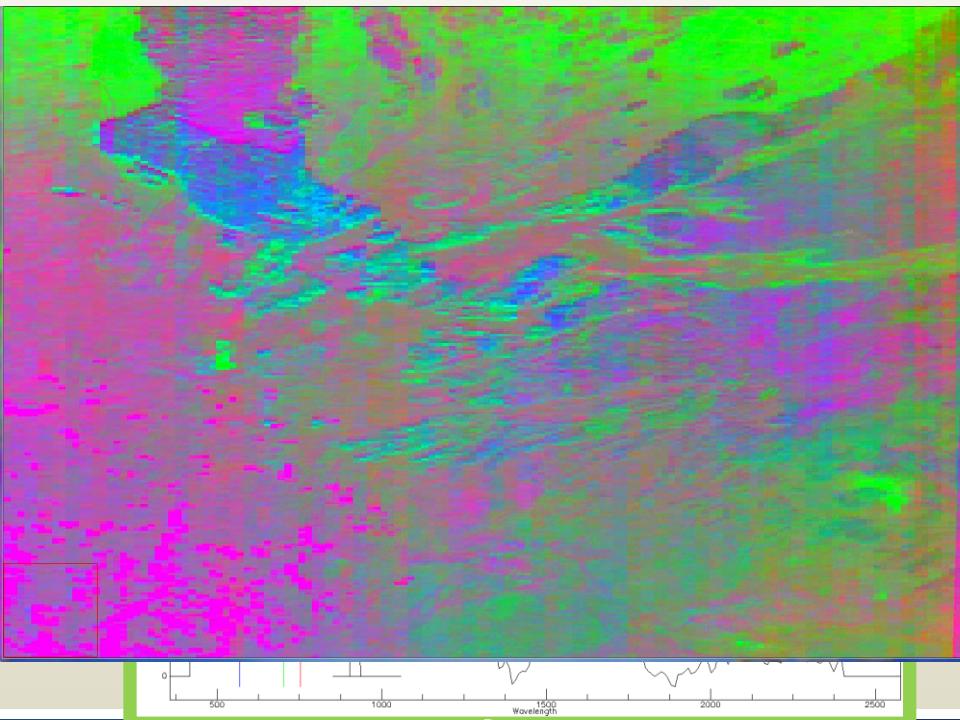
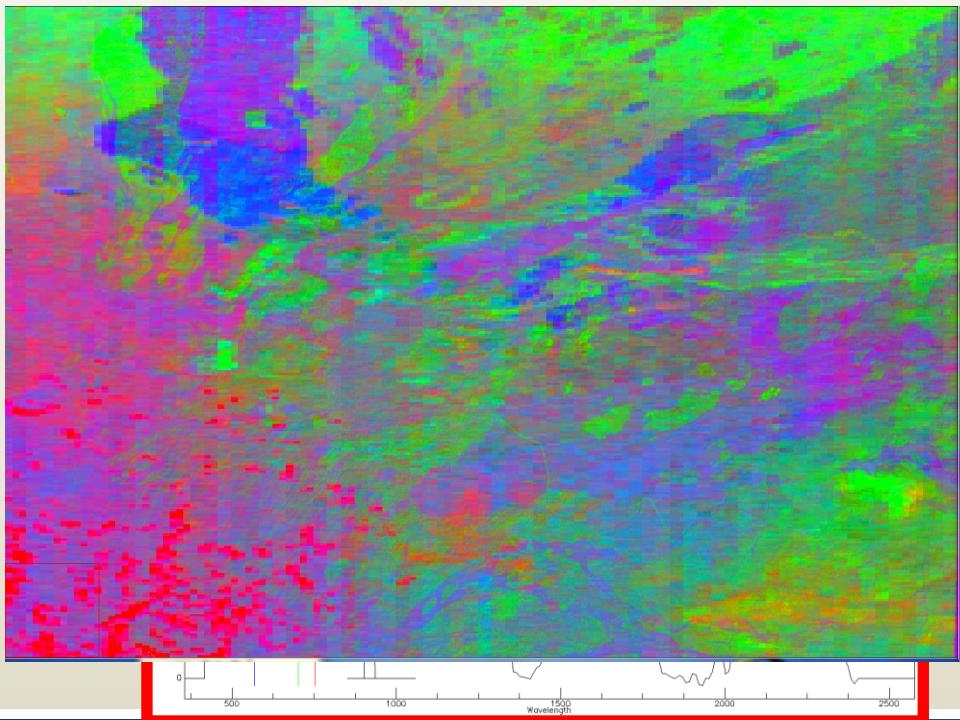
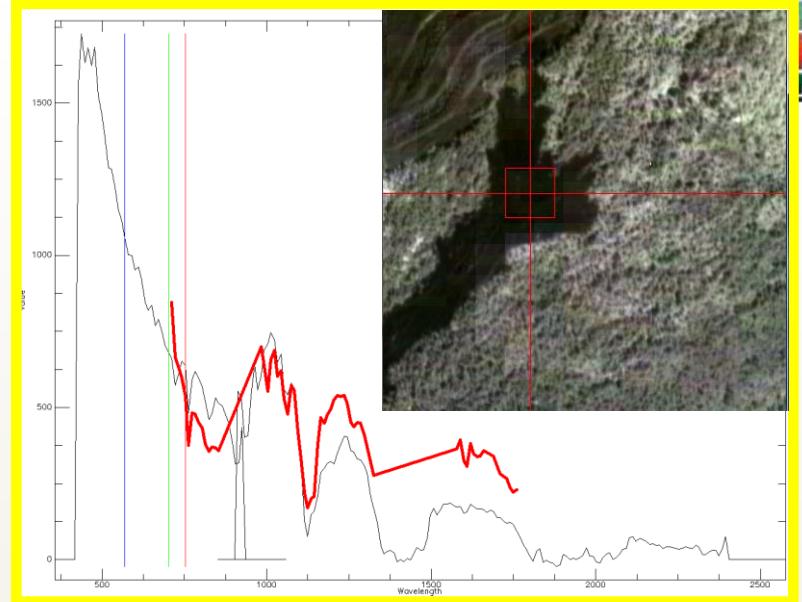
Gram Schmidt

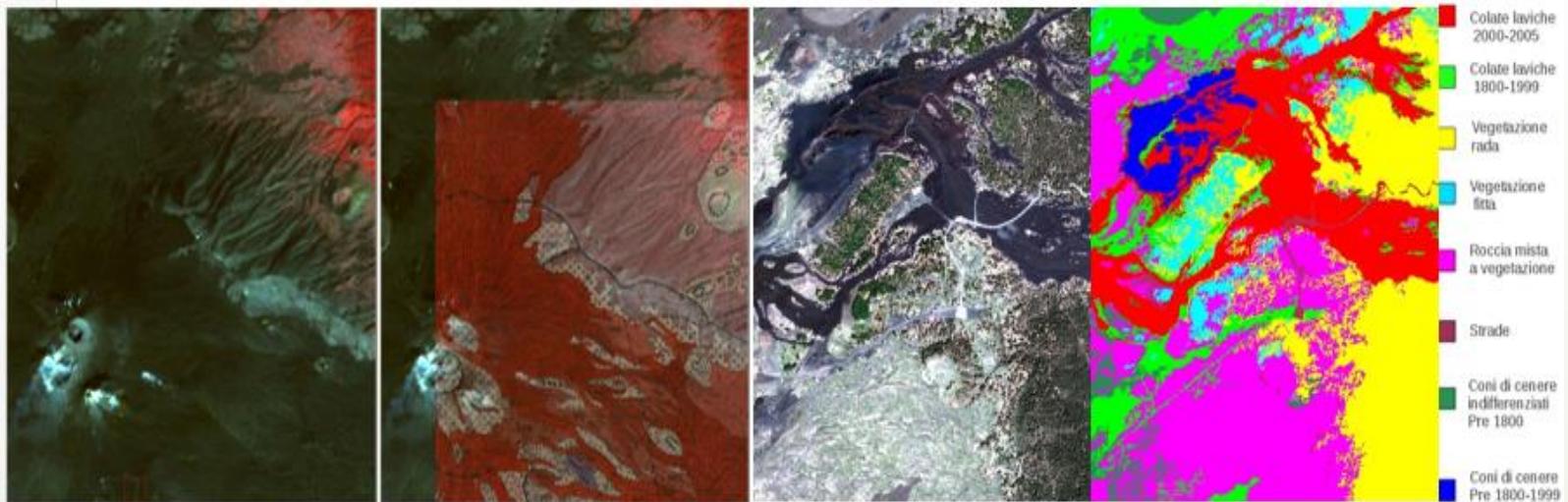
Altri

# VOLCANIC PRODUCTS CLASSIFICATION USING PAN SHARPENING TECHNIQUES

EO1-Hyperion, EO1-Hyperion+Quickbird, EO1-Hyperion (ridotto a 65 bande)+QB.

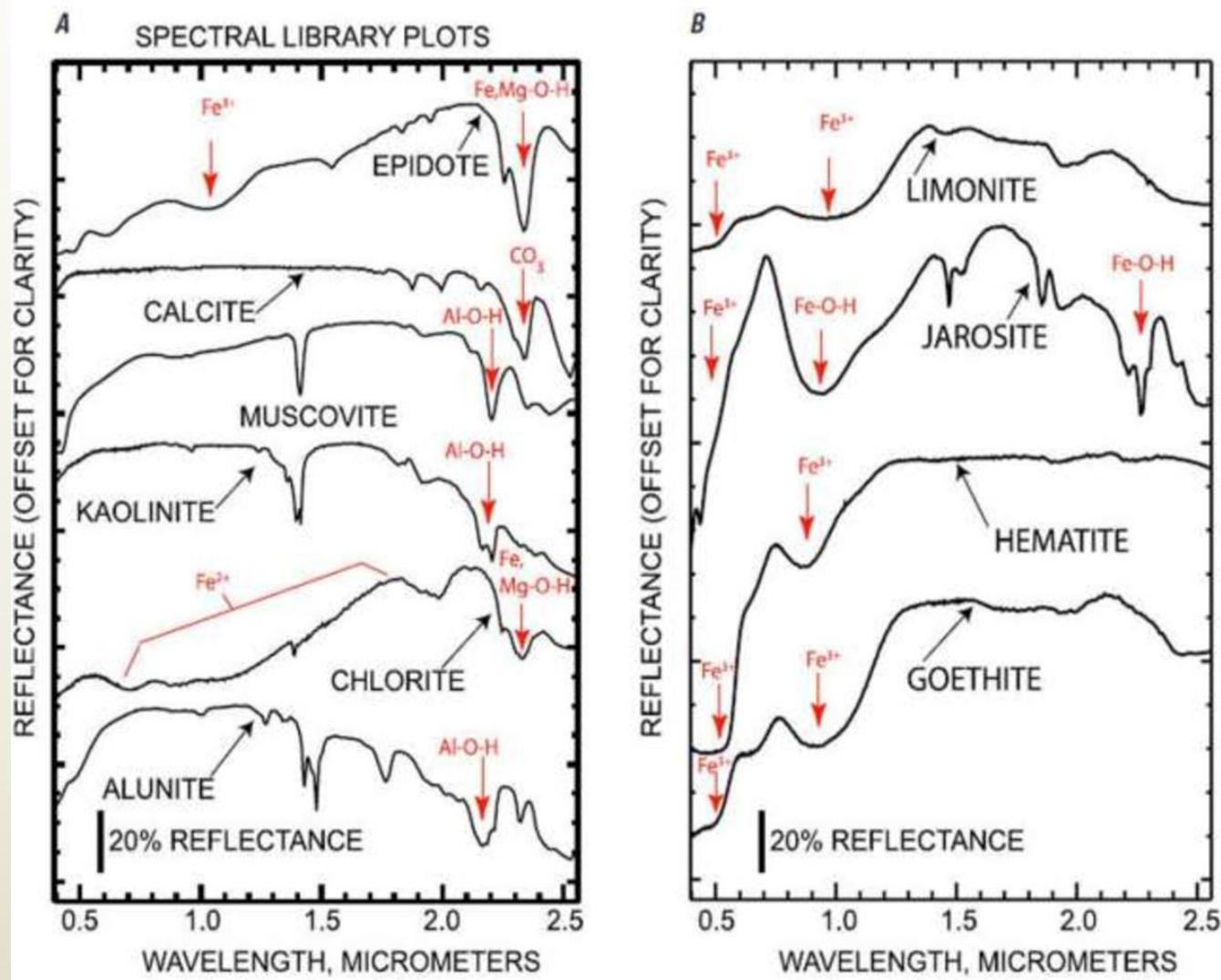




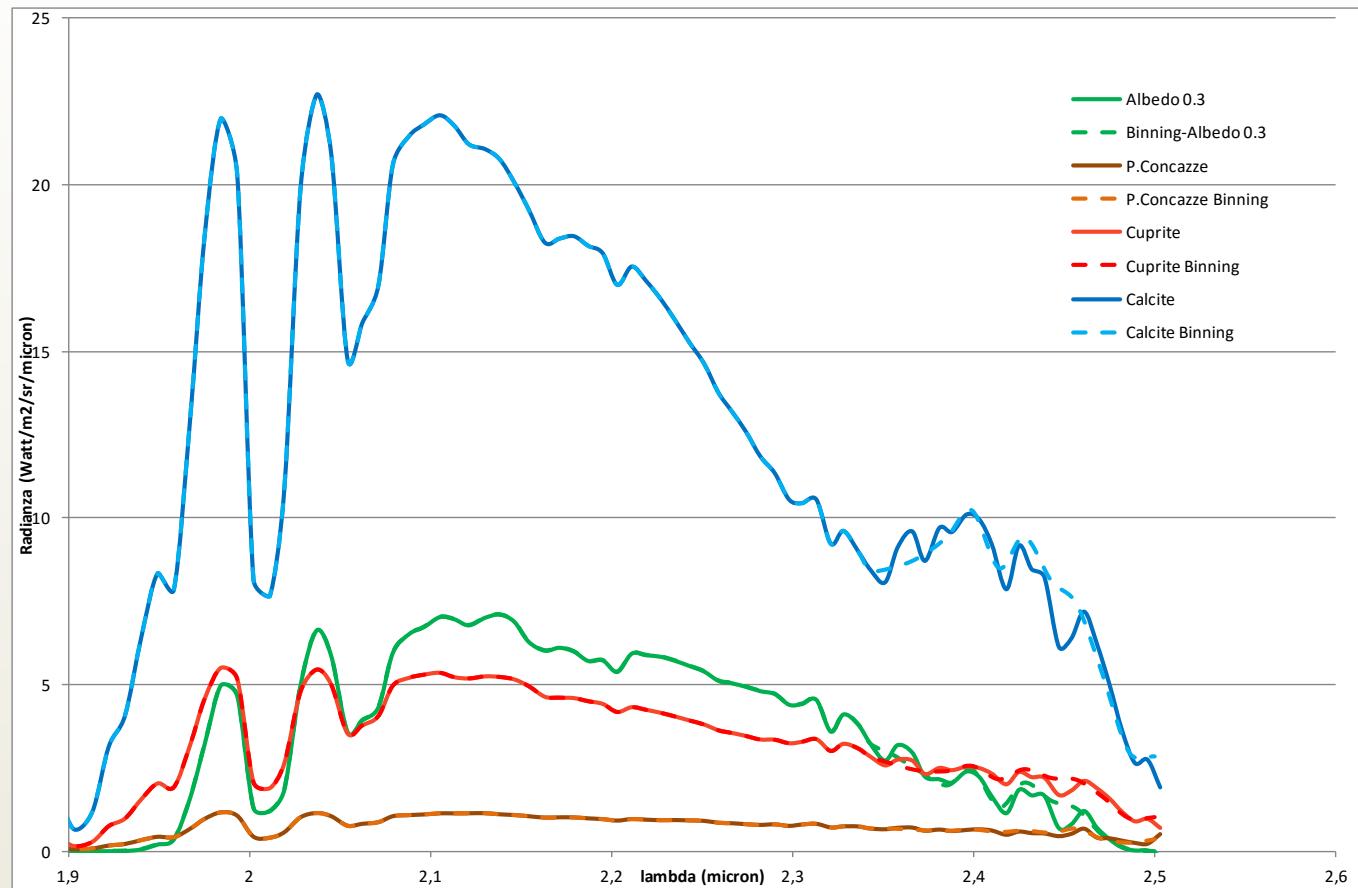


- The precision with which it obviously depends on the quality of remote sensing data and also used by the algorithm that will be used. The spatial resolution of the available sensors are matched by the spectral resolution whereas the characteristics of hyperspectral remote sensing is a viable source of alternative information and support to geological mapping, lithology and soil. Since the hyperspectral sensor observes an area of finite extension ( $30 \times 30$  m) that contains a variety of different materials, the spectrum of reflectance of the pixels in the remotely sensed images can be generally interpreted as a "mixture" of endmember spectra

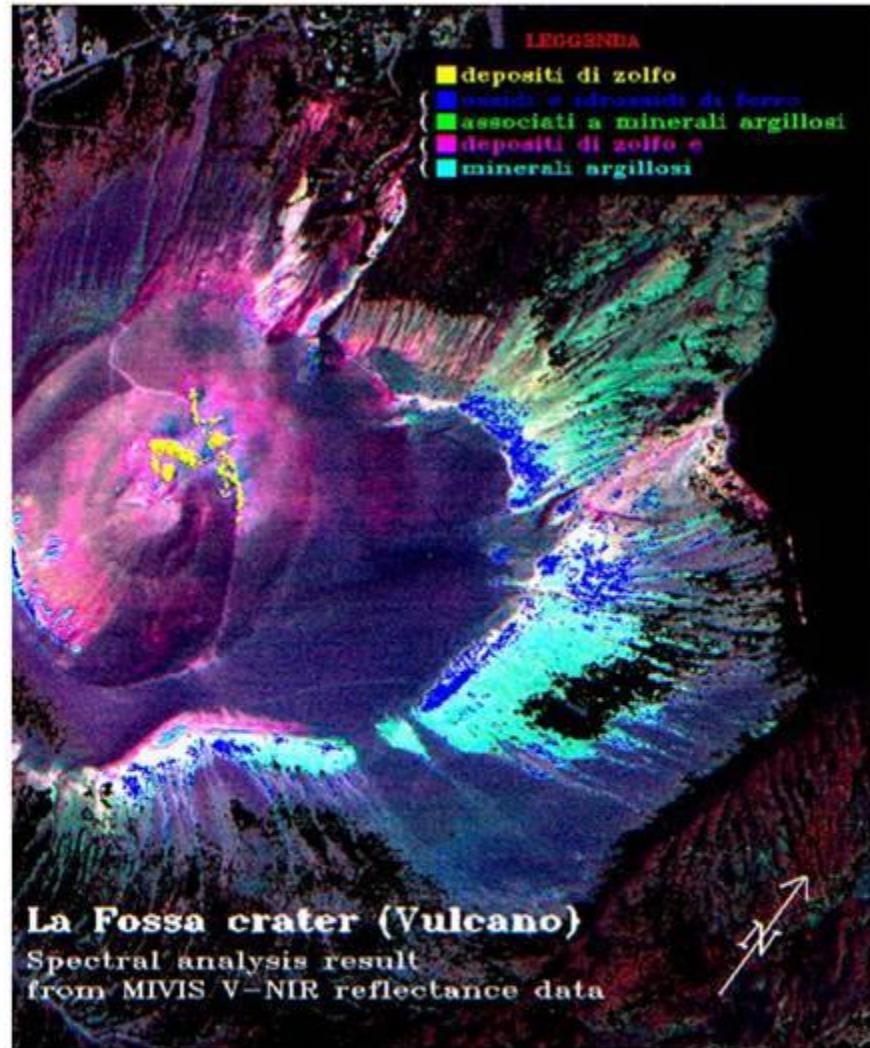
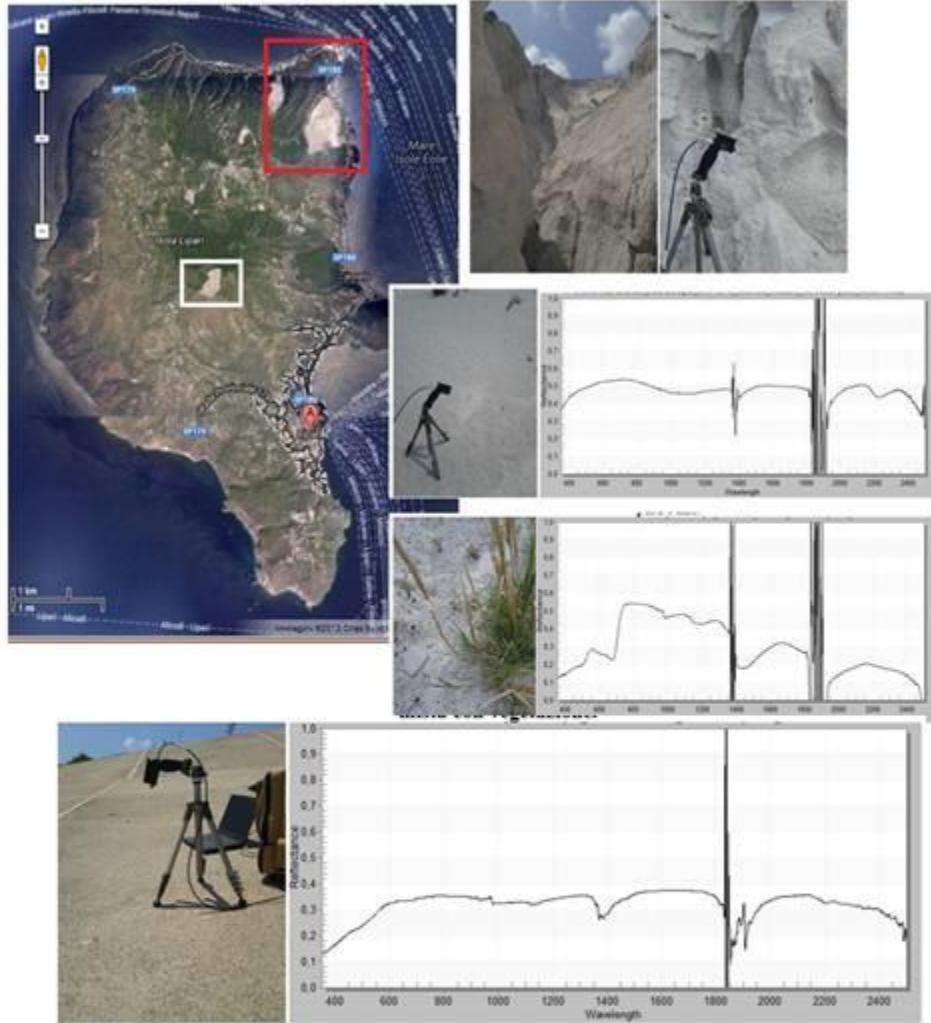
# MINERALOGY AND SOIL ANALYSIS



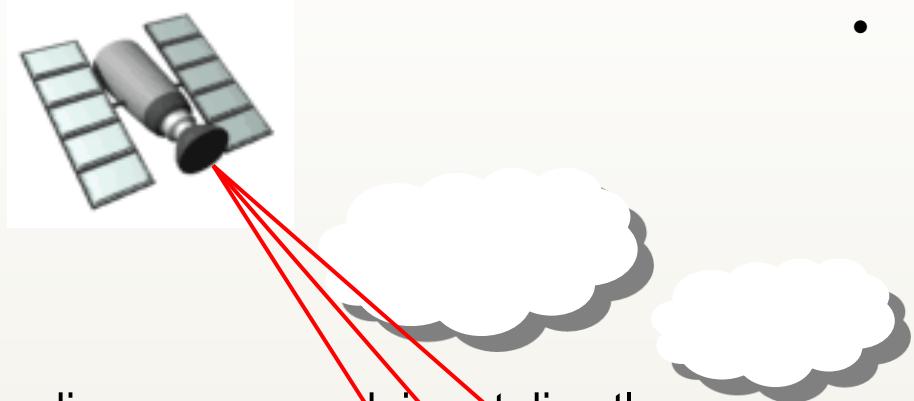
## SIMULATION OF RADIANCES FOR DIFFERENT MATERIALS AND TEMPERATURE



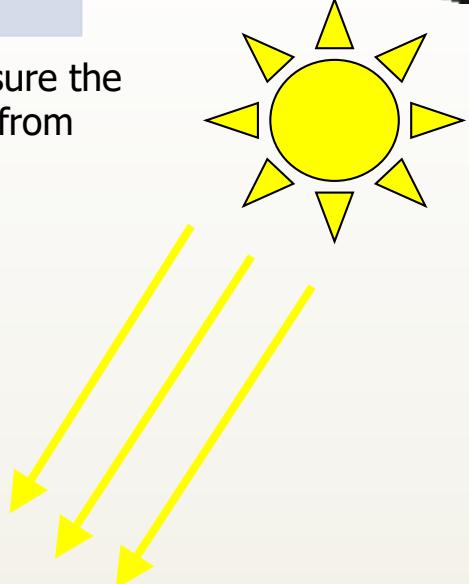
Radianza simulata PRISMA-like considerando come albedo: 0,3 (verde), calcite (azzurro), cuprite (rosso) e Piano delle Concazze (marrone). La linea continua rappresenta la radianza ottenuta considerando le 171 bande nello SWIR, in tratteggiato la radianza ottenuta considerando il binning sulle ultime bande dello SWIR (e quindi su 160 bande)



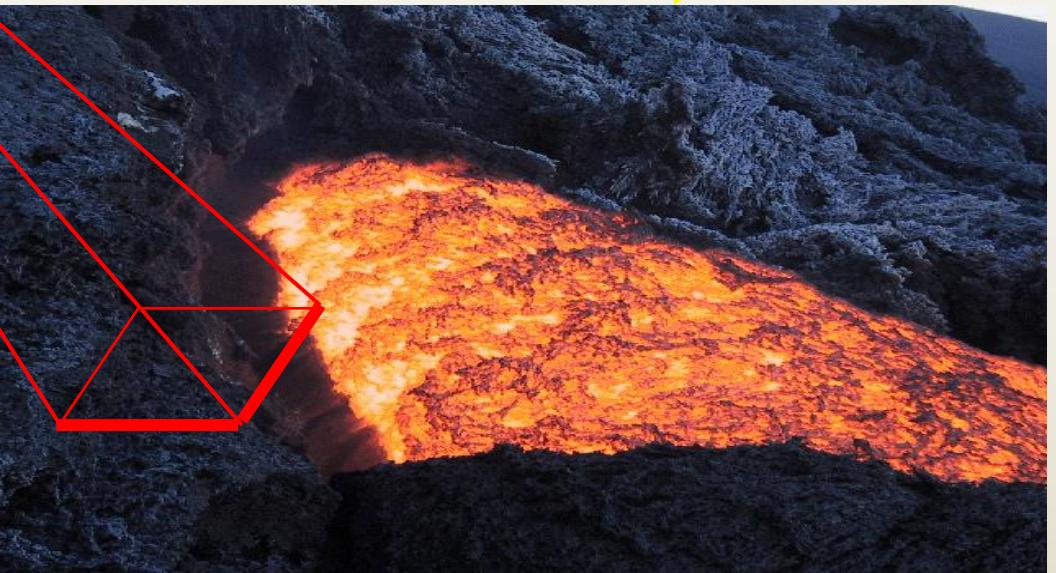
# Lava Thermal analysis



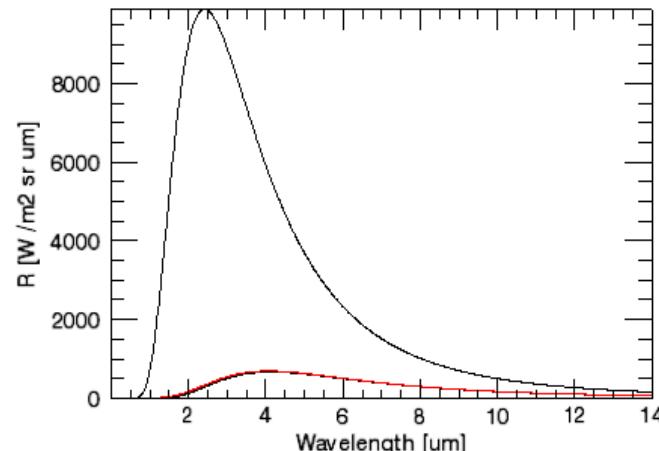
- The sensors measure the spectral radiance from the Earth



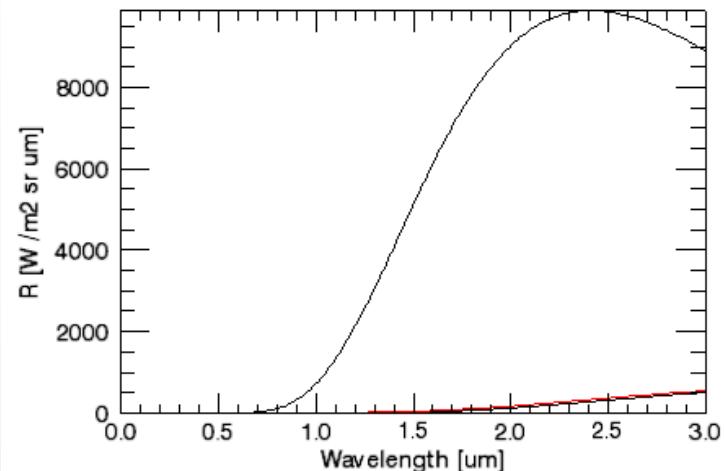
- The radiance measured is not directly related to the surface temperature but depends on:
  - The surface emissivity
  - Atmospheric absorption
  - The solar radiation



Dr. Valerio Lombardo



Radianza teorica tra 0.5 e 14 micron

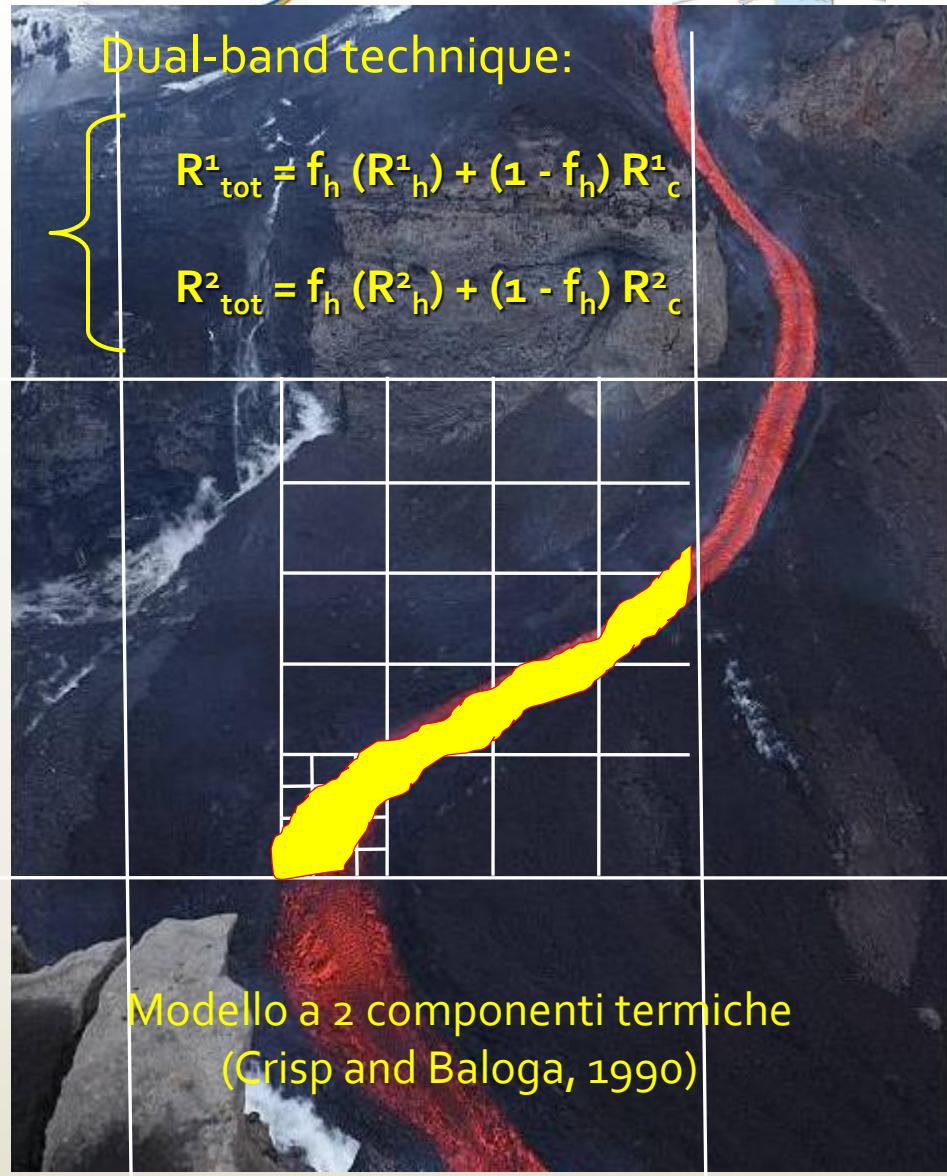


Radianza teorica tra 0.5 e 3 micron

In the case of radiant pixels occupied by lava, the emission peak (Planck curve) moves towards the shorter wavelengths as the temperature increases. At the spatial resolution of PRISMA pixels can be considered to be occupied by a predominant fraction of the crust at a temperature  $T_c$  and a complementary fraction of molten lava that radiates at higher temperature  $Th$ . Since the crust occupies a portion much larger than the one occupied by the molten lava, the integrated curve (red curve in Figures) corresponds to  $T_c$  of crust. Using a temperature of the molten part ( $Th$ ) of  $1200$  ° K and a temperature of the crust ( $T_c$ ) at  $700$  ° K is obtained an integrated theoretical radiance

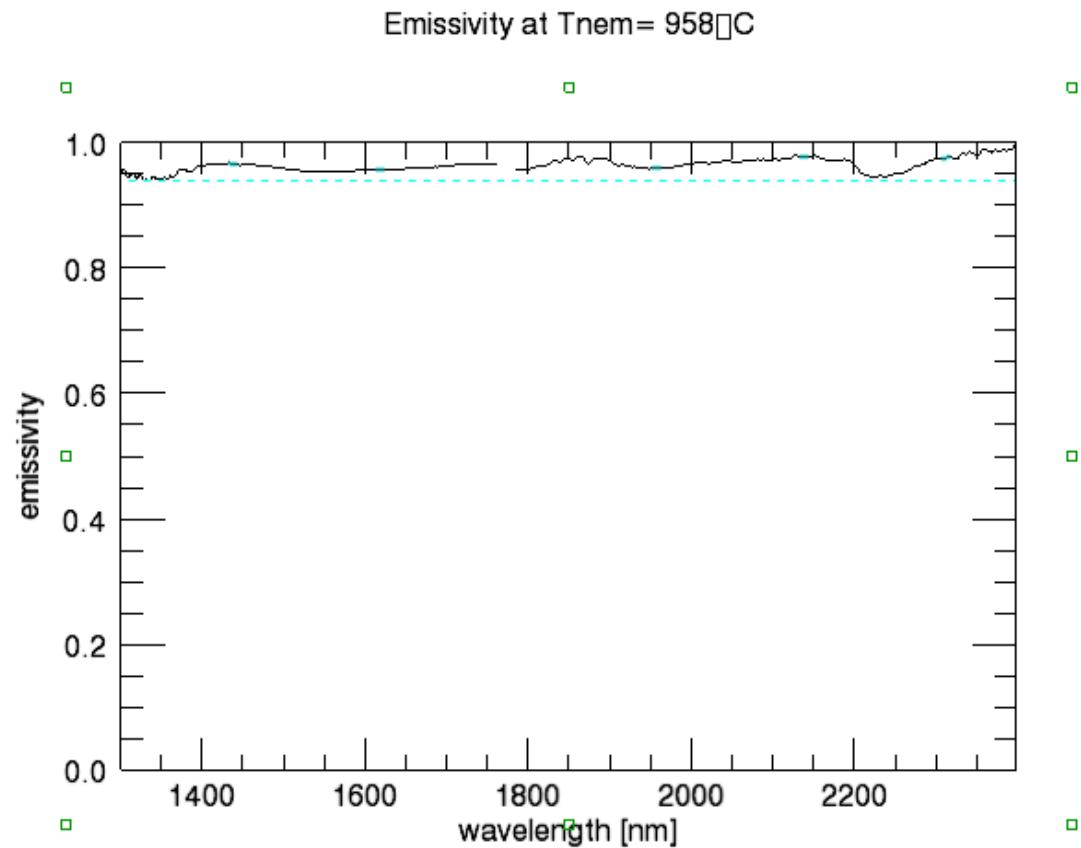
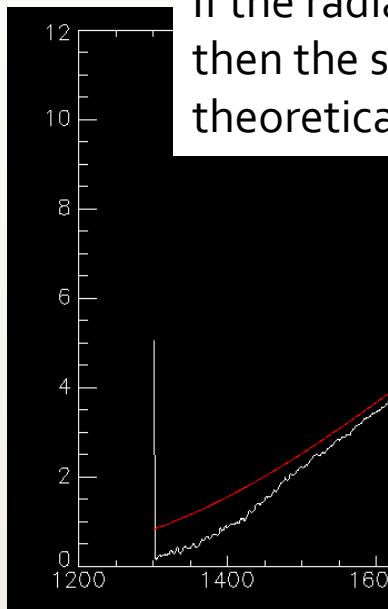
The temperature of the lava flows at pixel scale is never homogeneous but there are always at least two separate thermal components:  
Th (hot)  
Tc (cold)

Occupying a fraction of the pixels ( $f_h$ ).  
So it has been developed a new algorithm to determine simultaneously the spectral emissivity and the various thermal components using the spectral and radiometric characteristics provided for PRISMA



# Temperature/Emissivity Separation

If the radiance has been corrected for the solar and atmospheric contribution then the spectral emissivity is the ratio between the experimental and theoretical curve curve (black body):



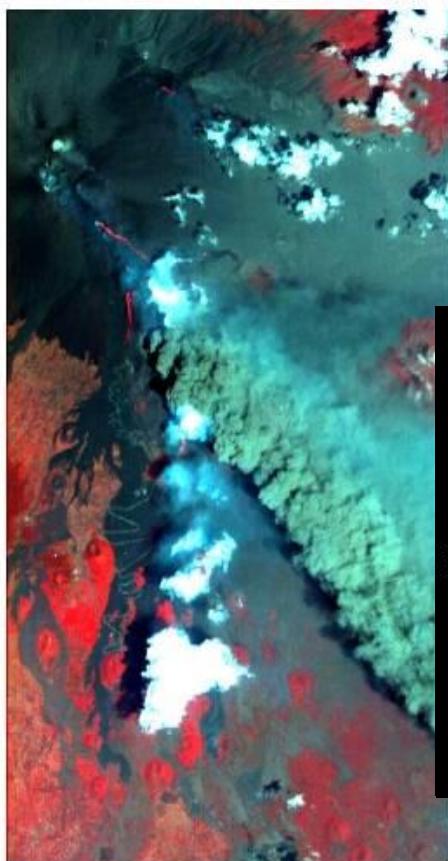
# Algoritmo "Draping"

For this reason, a new algorithm called "draping" was implemented. The "draping" algorithm allows the simultaneous calculation of Th, Tc, fh, and spectral through a series of iterations between the radiance of the experimental curve and a set of theoretical curves calculated by varying Th, Tc, in the previous equation fh imposing Rtot always equal to Rmax at the wavelength I<sub>max</sub> wave.

Requirements for application of the algorithm:

- Ispettrale SWIR (PRISMA)
- High radiometric resolution
- Radiometric data unsaturated

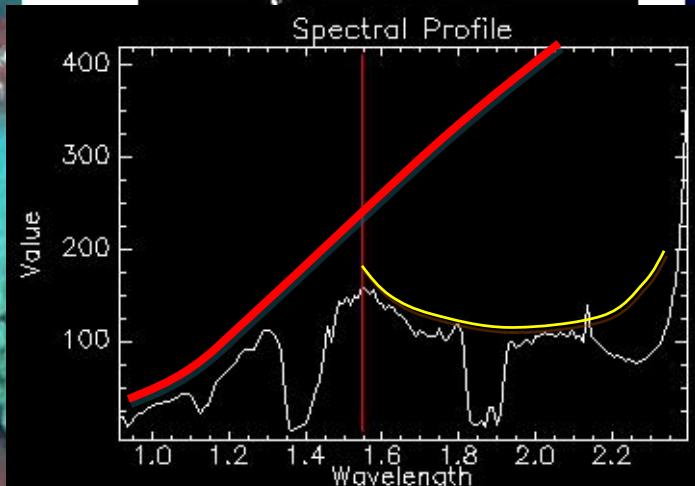
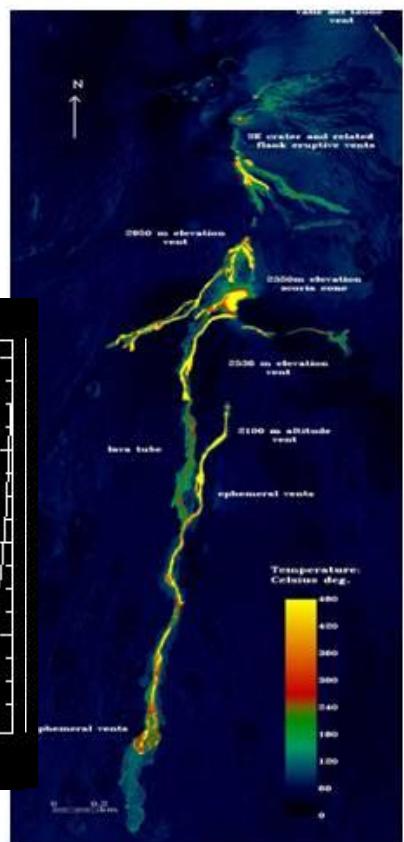
EO-1 Hyperion



Hot spot detection

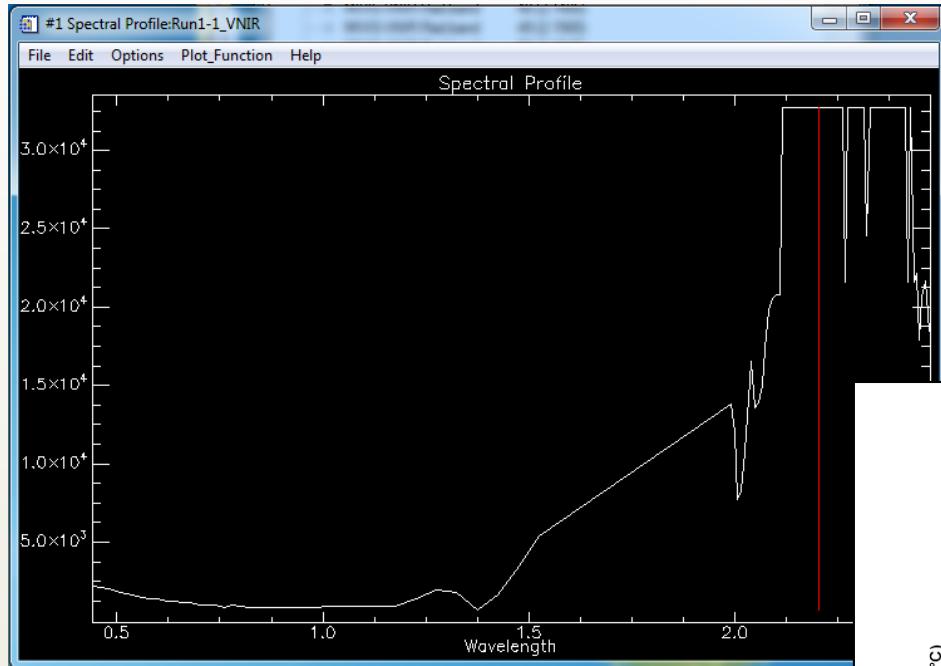


Temperature retrieval



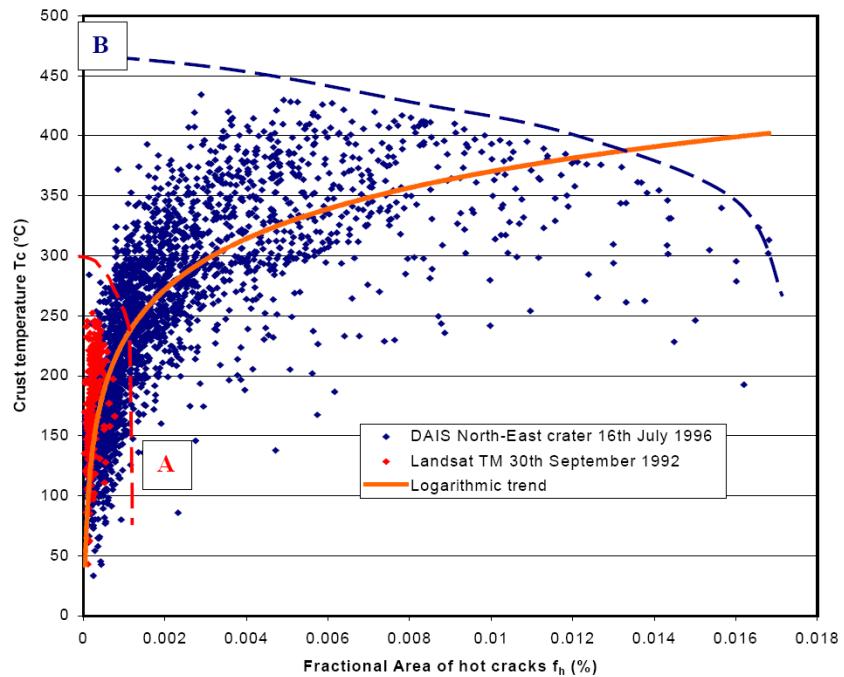
Hot-Spot Detection (Hyperion and MIVIS) and temperature retrieval during the Etna eruption (2001)

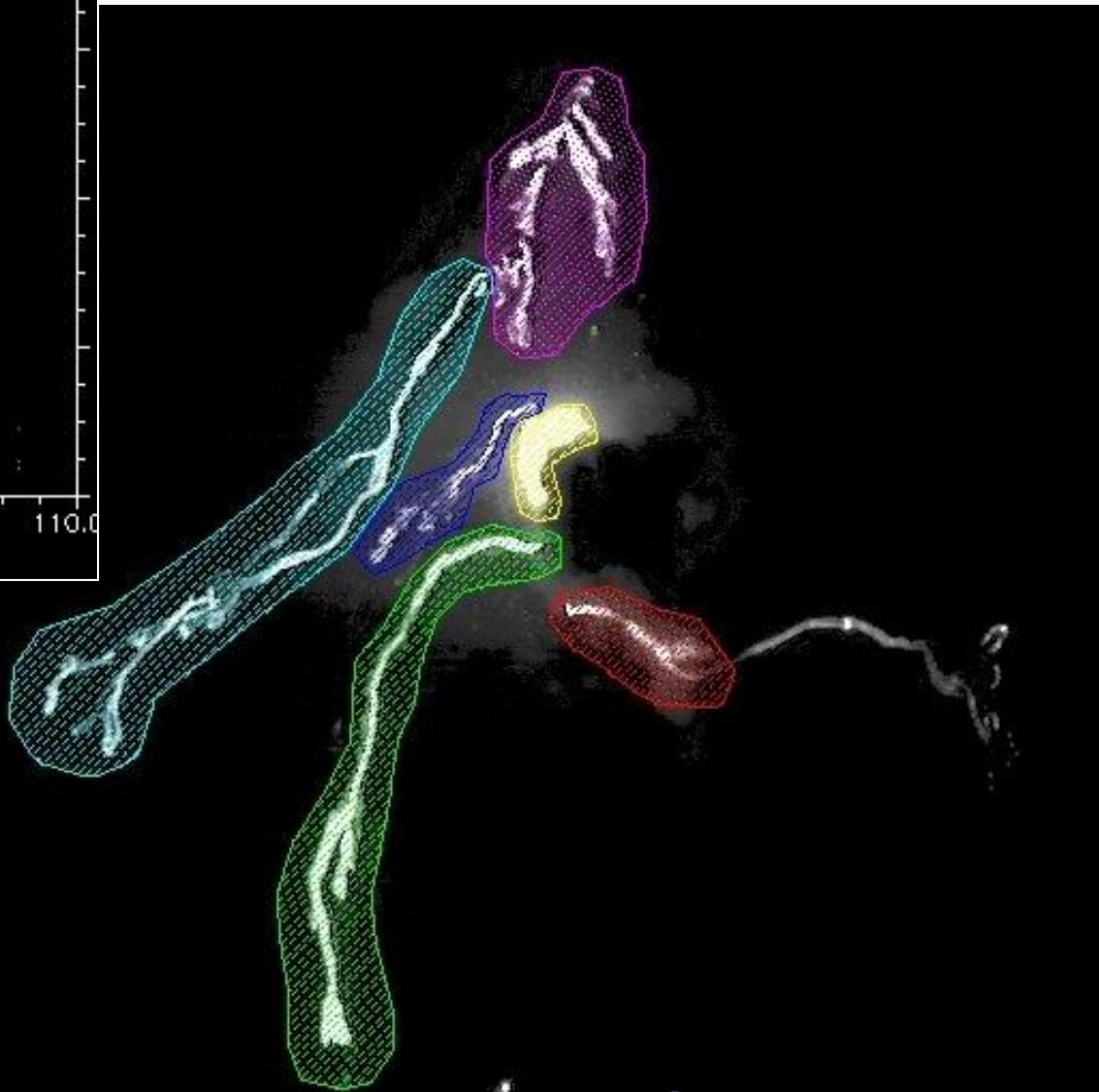
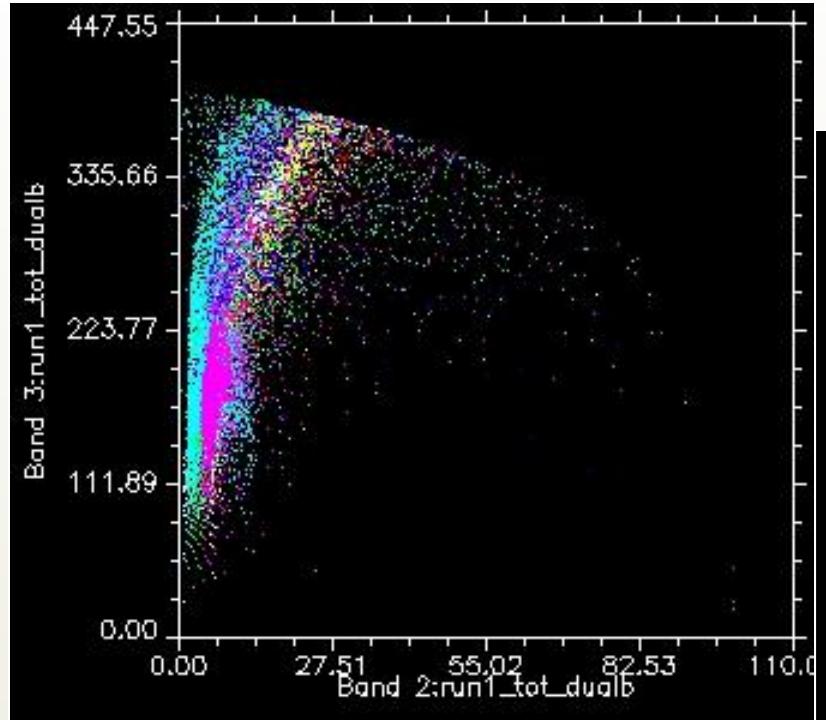
# Applicazione su dati aerei



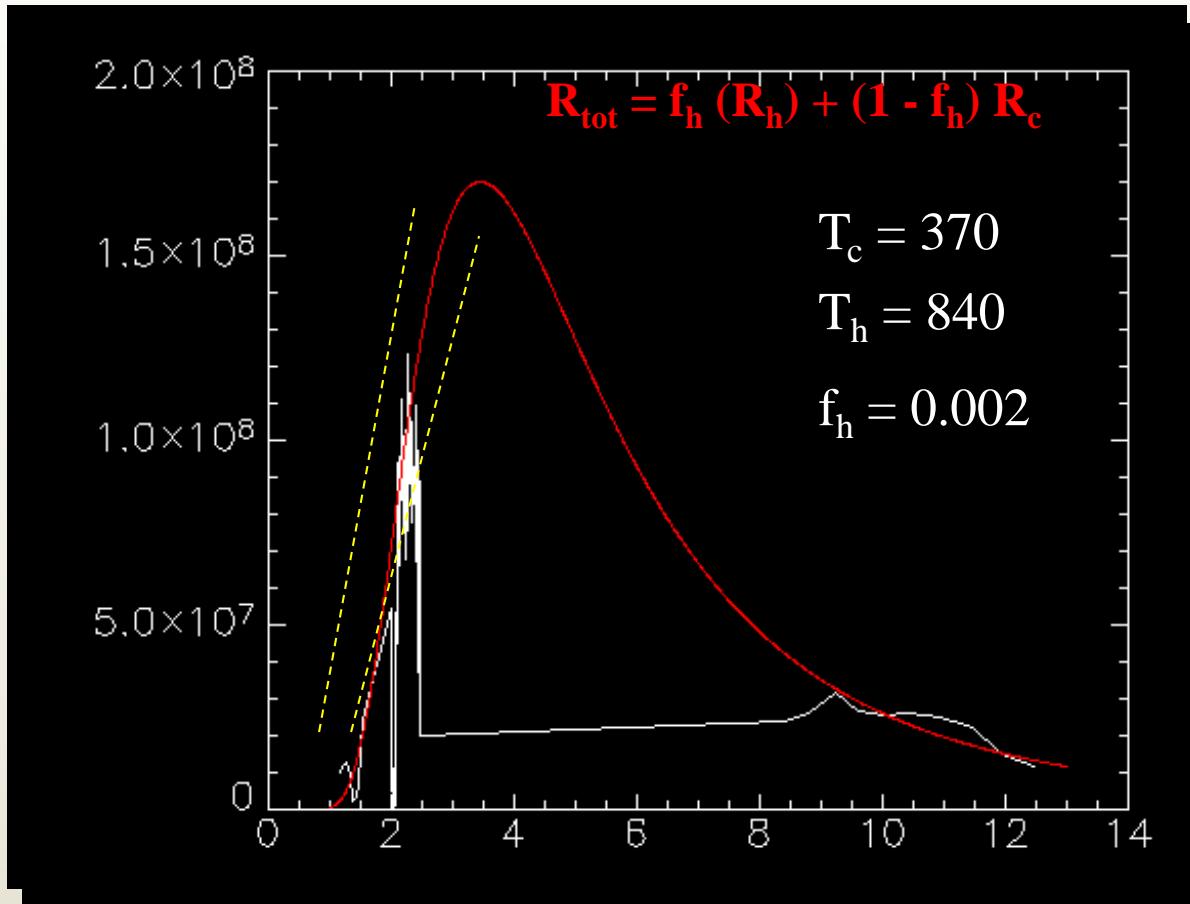
I dati aerei come il MIVIS permettono il settaggio del *gain* e *offset* e quindi di innalzare il limite di saturazione

Confronto tra soluzione  $T_c$  vs  $f_h$  ottenute con dati DAIS e dati ASTER.



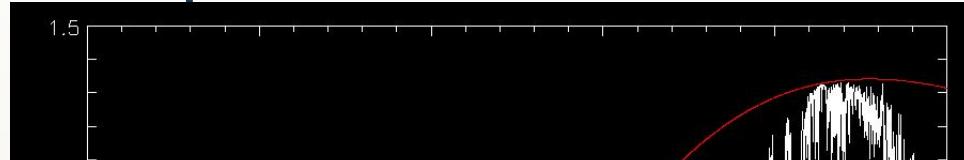


## VIS-SWIR channels are not enough to constrain the planck curve at specific temperature



Integrated curve: theoretical curve (red) and real spectrum as detected by MIVIS sensor (white).

# Requisiti del sensore

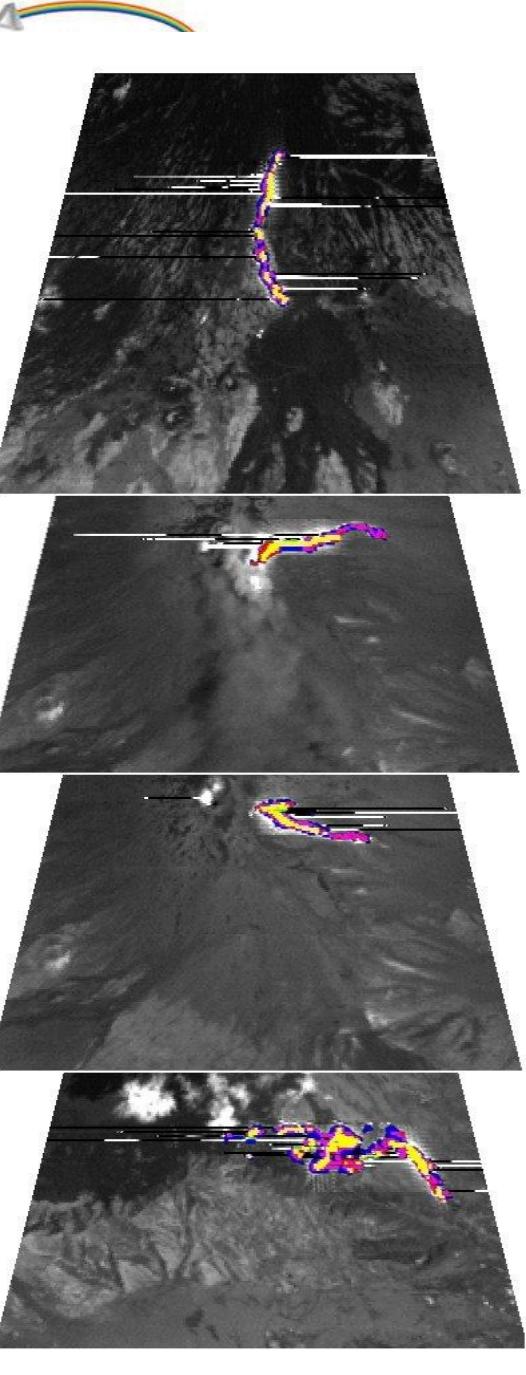
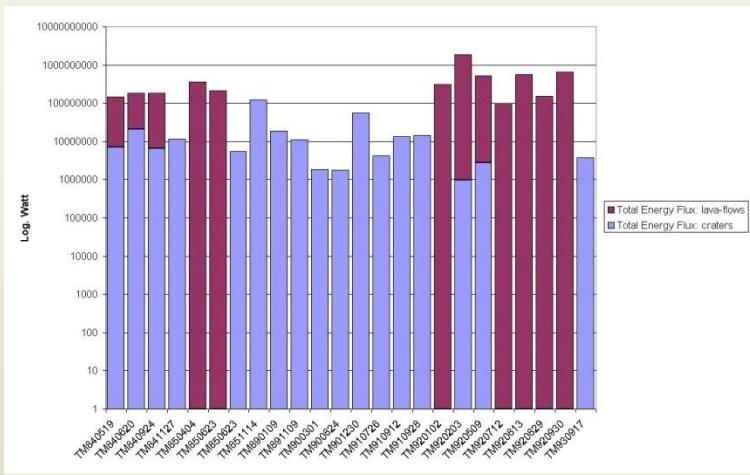
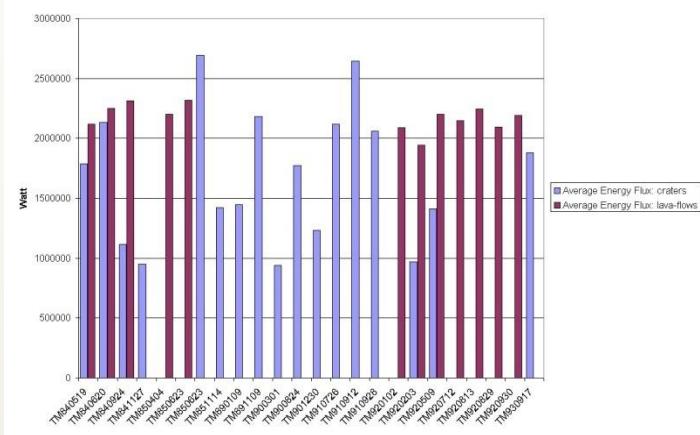


Simulazione MODTRAN



Simulazione PRISMA inserendo spettro di radianza a 1000°C. Si è notata solo una leggera saturazione.

# Temporal Comparison of Thermal and Geometrical Features



# PRISMA PRODUCTS

Starting from the radiance of SWIR bands with the algorithm of "Draping" will be able to get from the data PRISMA the spectral emissivity and the sub-pixels temperatures.

Number of SWIR bands envisaged for PRISMA would apply this algorithm

Along with this requirement a dynamic range, however, it must be provided which prevents the saturation of the signal in the presence of high temperatures (as happens for Hyperion data).

A saturation value of 1000 °C would be ideal to ensure the observability of volcanic phenomena and fires. 500 °C are the theoretical limit of applicability of the algorithm.

# CONCLUSION

- PRISMA DATA COULD EFFECTIVELY SUPPORT VOLCANIC RISKS MANAGEMENT
- DATA FUSION WITH TIR DATA AND SAR DATA COULD BE OF HIGH INTEREST
- ATMOSPHERIC AND TOPOGRAPHIC CORRECTIONS ARE VERY IMPORTANT TO PRODUCE HIGH QUALITY DATA
- HIGH TEMPERATURE OBSERVATION ARE IMPORTANT BUT REQUIRES NON SATURATED PIXELS



**remote sensing**

IMPACT  
FACTOR  
**3,036**

Special Issue

**Widespread Applications Based on Hyperspectral Technologies from Space**

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Deadline for manuscript  
submissions:

30 June 2018

**Message from the Guest Editors**

Dear Colleagues,

The aim of this Special Issue is to highlight the impact of the past hyperspectral missions and foresee the effectiveness of the future ones. We would like to invite submissions on the following topics:

- Integration and comparison of new hyperspectral image data/constellation;
- Natural processes and human activities and their interactions, including archaeology;
- Environmental and natural hazards and risks reduction;
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